

*Contributed Talk //*

Gamma-Ray Bursts (GRBs)



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Since the first detection of gamma-ray bursts (GRBs) in 1967, GRBs have been an active subject of study with many questions still left unanswered. Though the physics of GRB afterglows is relatively simple and more well-known, there is still much discussion regarding the prompt emission phase of GRBs. Several models, including synchrotron models with both ordered and random magnetic fields, synchrotron self-Compton models as well as Compton-drag models and photospheric models have been proposed in order to explain the Band-like spectra of GRB prompt emissions. Seeing as both synchrotron radiation and anisotropic Comptonization (as expected in both photospheric and Compton-drag models) may lead to X-ray polarization, polarization predictions from the various proposed models should inevitably be used in conjunction with X-ray polarization measurements by POLAR and possibly the future COSI missions in order to investigate the emission mechanisms of GRBs. In light of this, we develop a Monte Carlo polarization code, heavily inspired by the MAPPIEs code, to estimate Compton polarization signatures from various GRB models in order to constrain and hopefully discriminate between the available models.

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